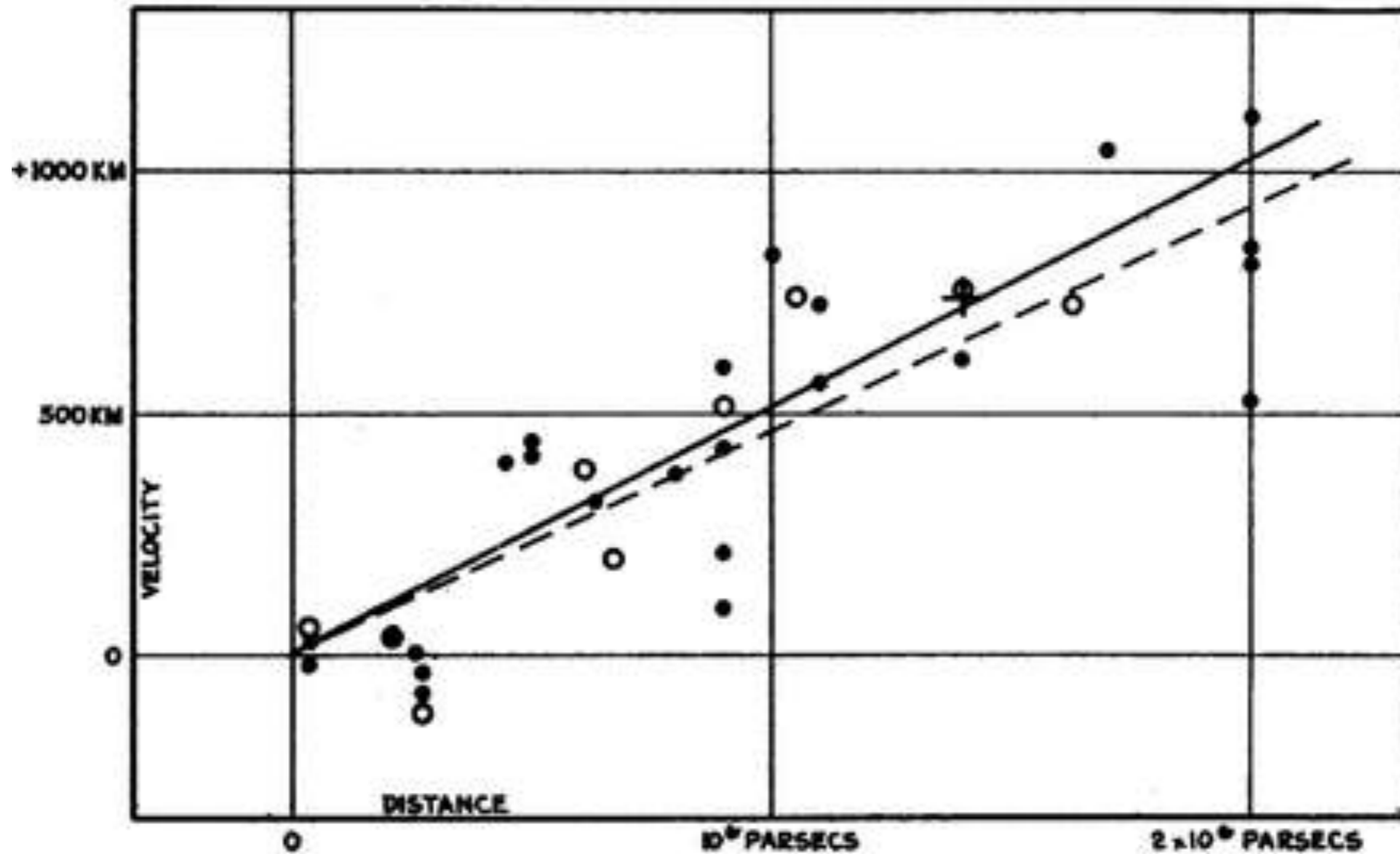


Hubble's original 1929 diagram

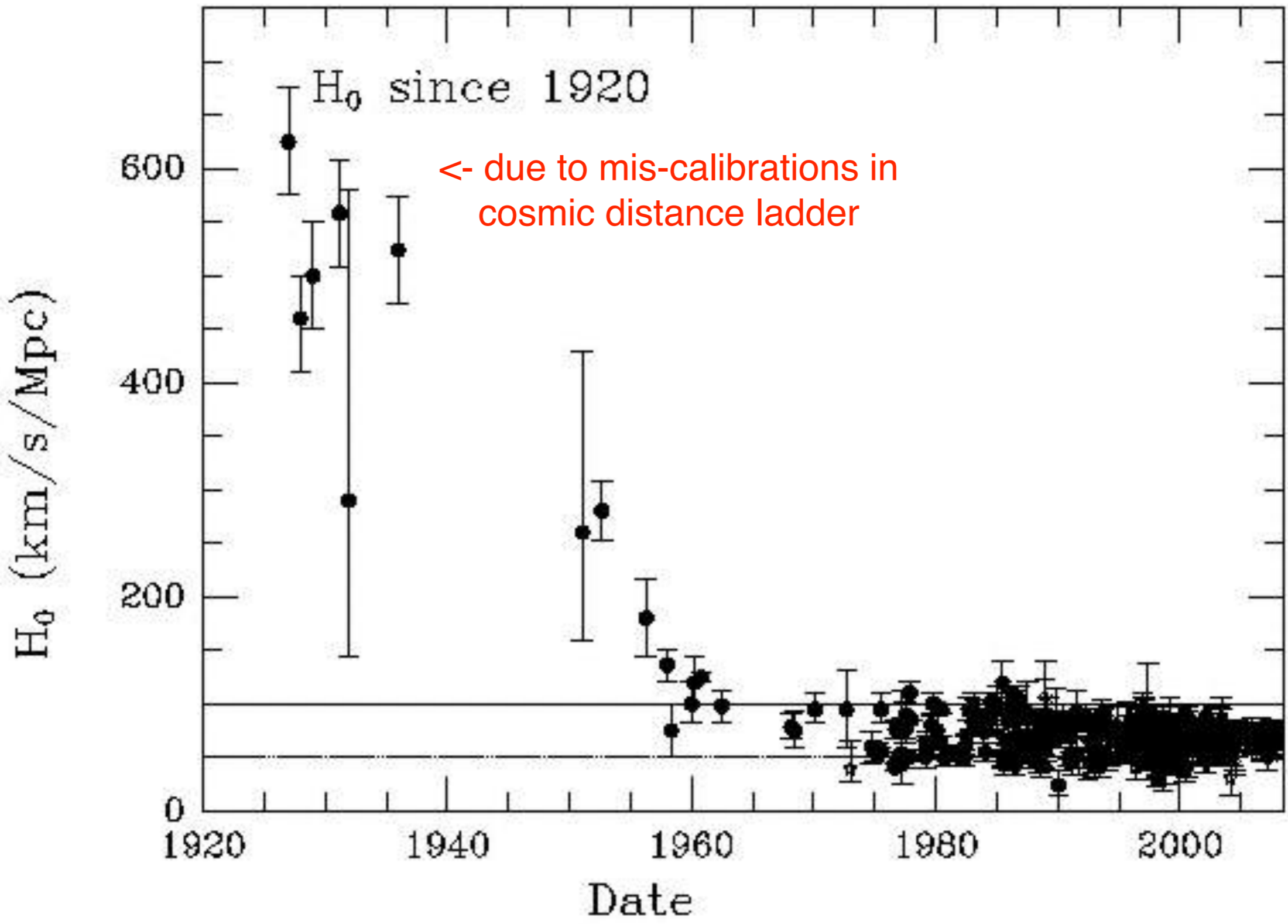
Velocity-Distance Relation among Extra-Galactic Nebulae.



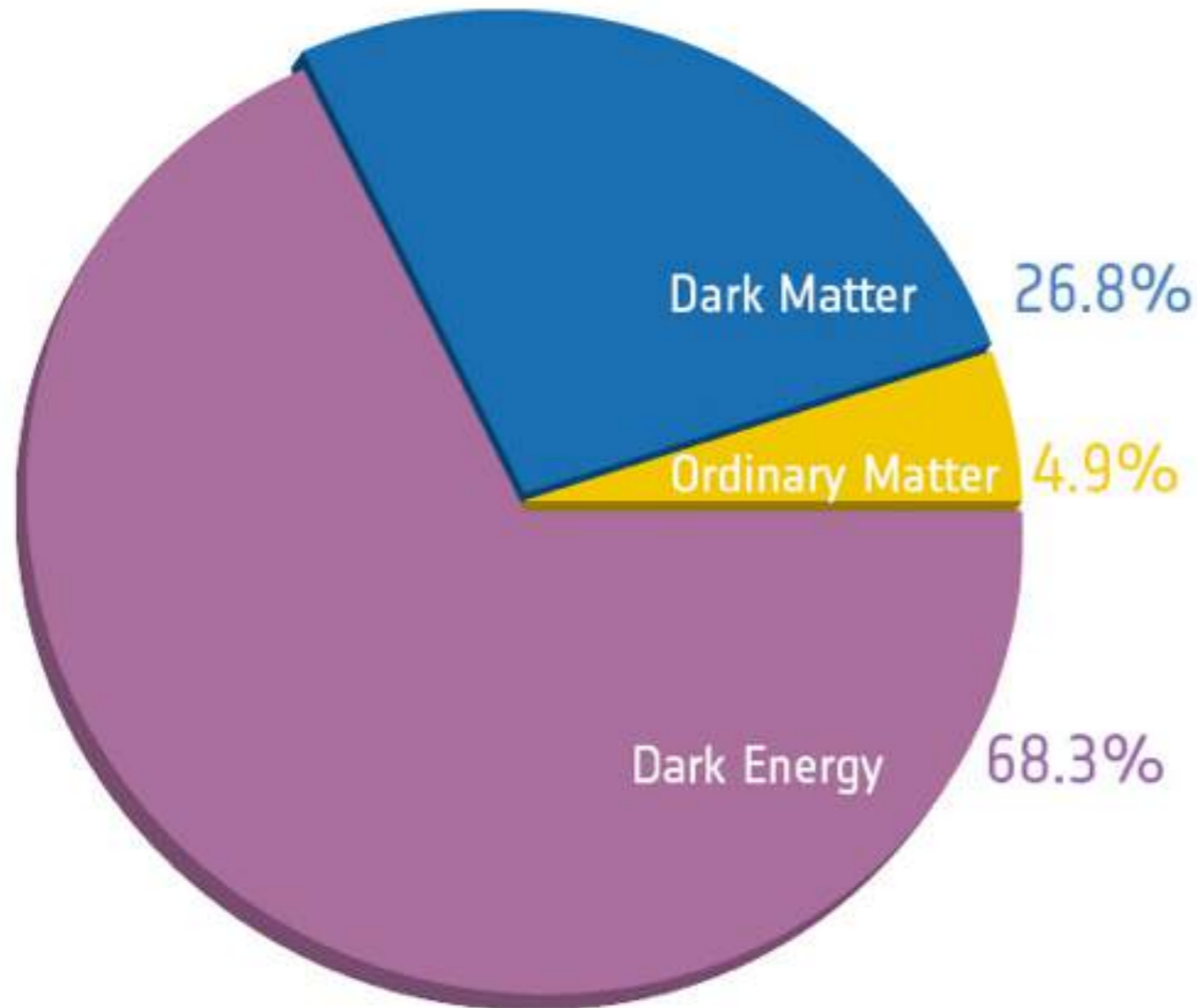
Distances determined using stars of known luminosity, including Cepheids

Note: velocity units should km/s

Measurements of the Hubble constant versus time



The standard Λ CDM cosmological model



In general:

$$\Omega_x \equiv \frac{\rho_x}{\rho_{\text{crit}}}$$

ρ_{crit} = critical density
to close universe

Our Universe:

$$\Omega_{\text{tot}} = 1$$

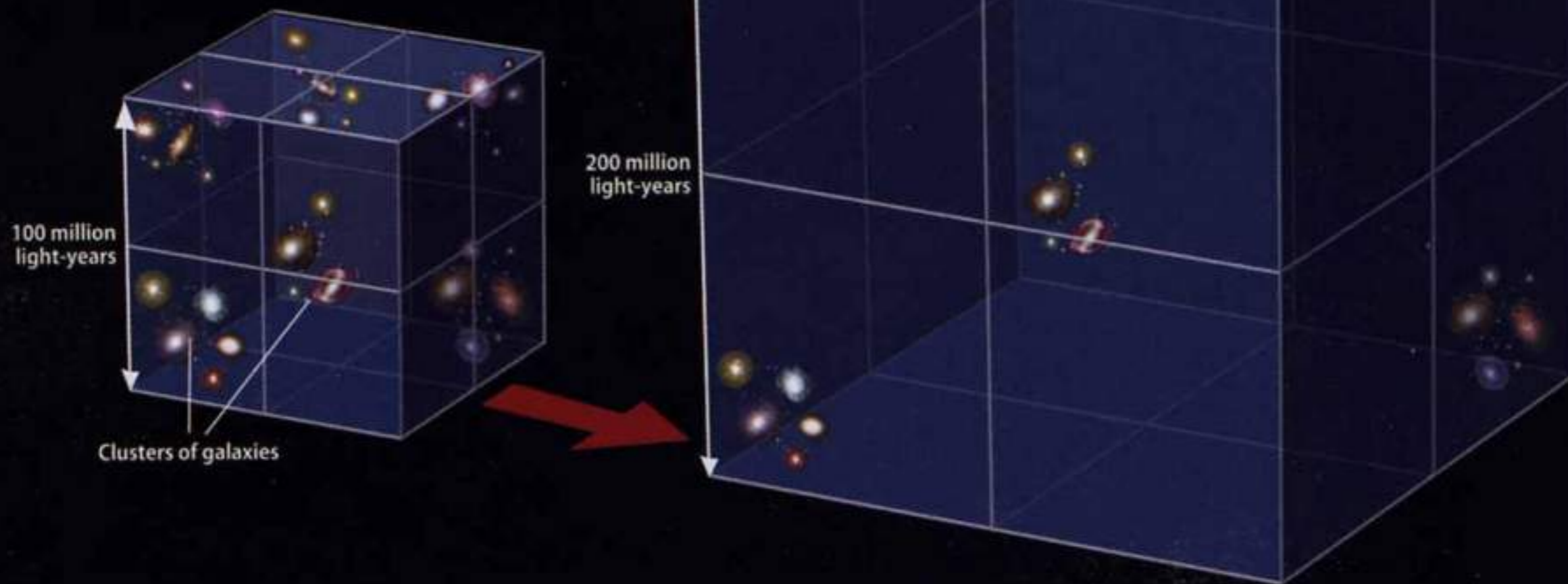
$$\Omega_{\text{DM}} = 0.27$$

$$\Omega_{\text{b}} = 0.05$$

$$\Omega_{\text{M}} = \Omega_{\text{DM}} + \Omega_{\text{b}} = 0.32$$

$$\Omega_{\Lambda} = 0.68$$

Cosmic expansion



The Nobel Prize in Physics 2015



Ill. N. Elmehed. © Nobel Media AB 2015.

Takaaki Kajita

Prize share: 1/2



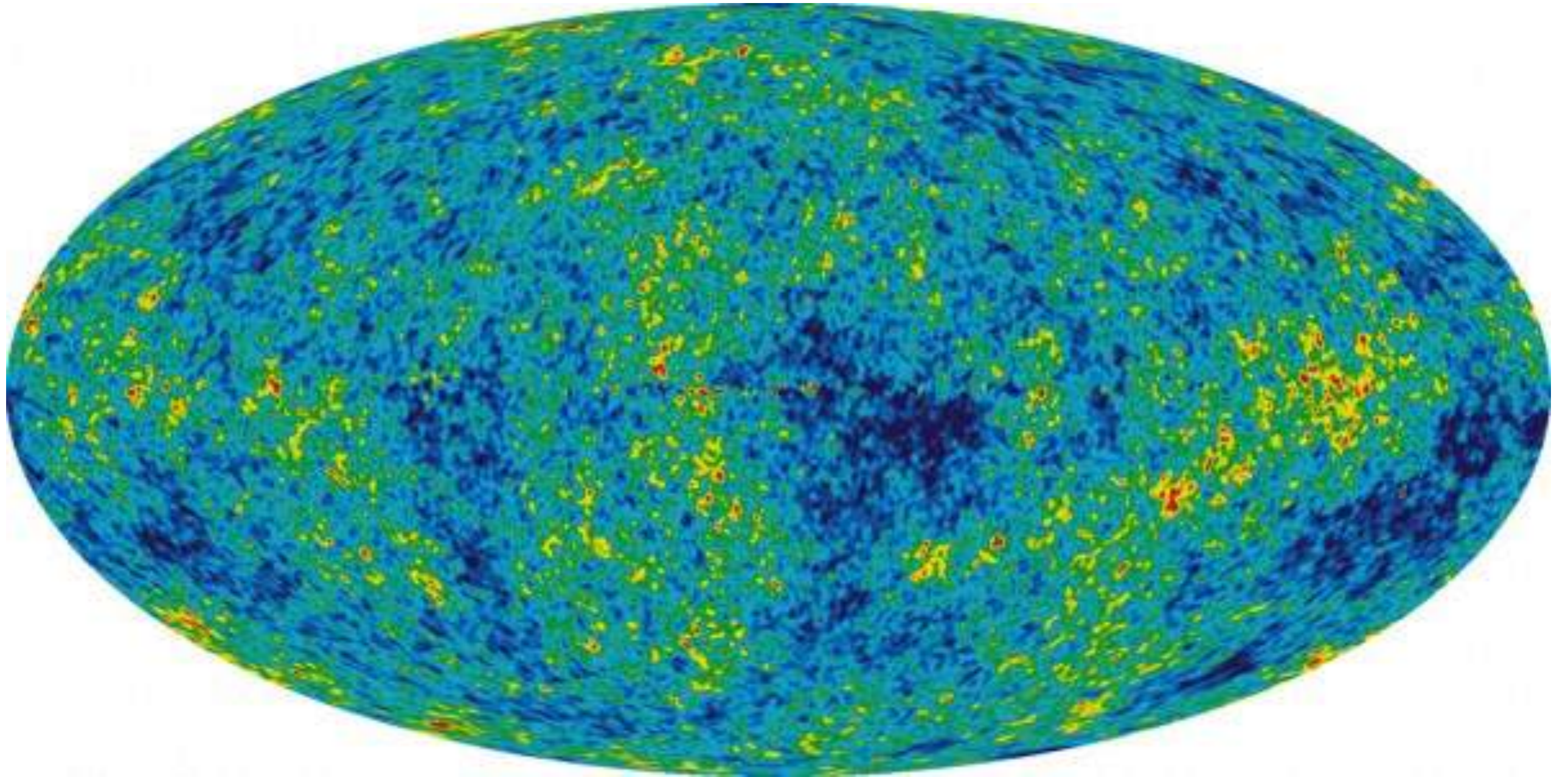
Ill. N. Elmehed. © Nobel Media AB 2015.

**Arthur B.
McDonald**

Prize share: 1/2

The Nobel Prize in Physics 2015 was awarded jointly to Takaaki Kajita and Arthur B. McDonald *"for the discovery of neutrino oscillations, which shows that neutrinos have mass"*

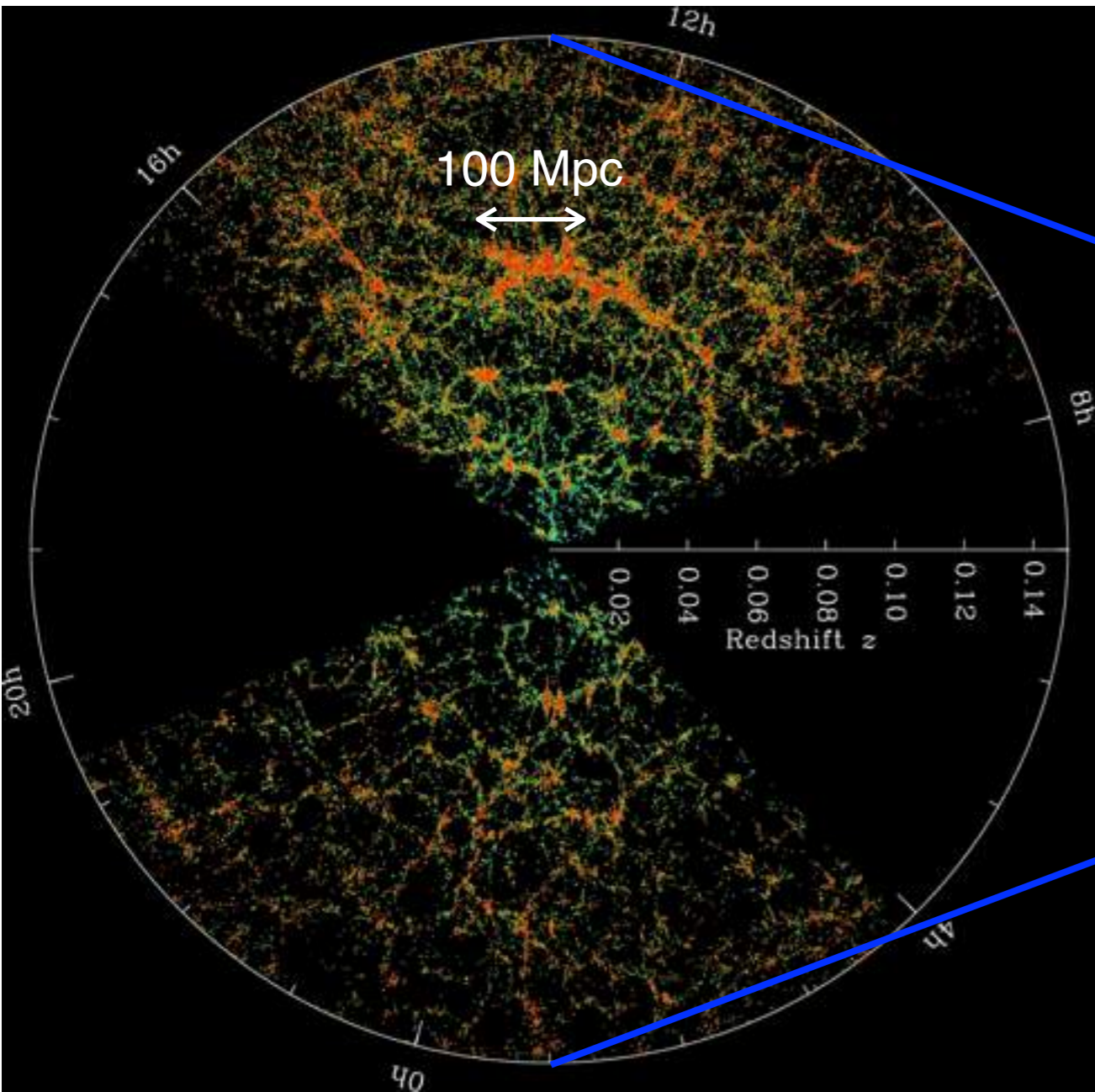
The CMB is uniform to 1 part in $\sim 10^5$



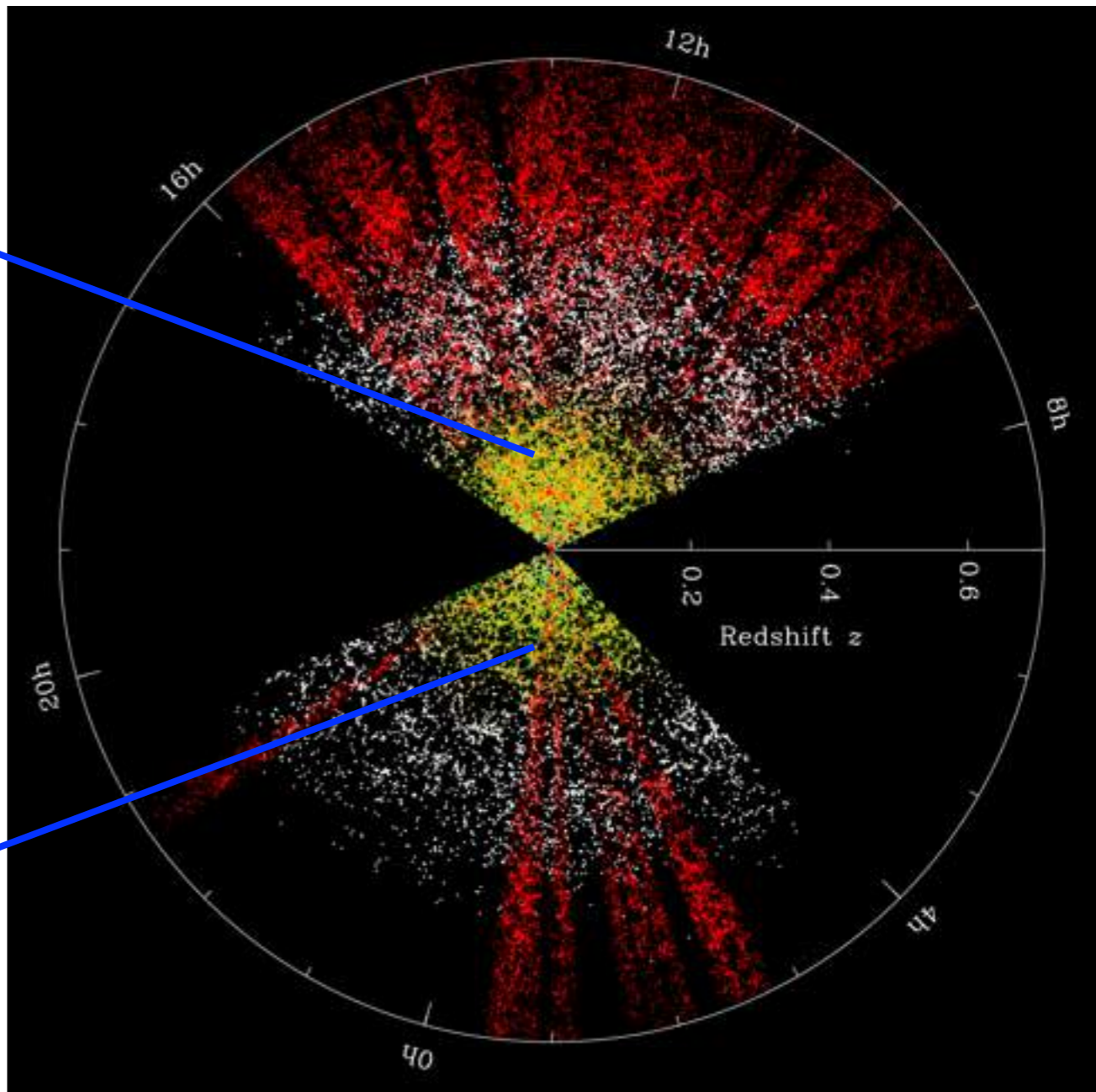
\sim density fluctuations at $z \sim 1,100$

Galaxies map the late time distribution of matter

Sloan Digital Sky Survey galaxy distribution



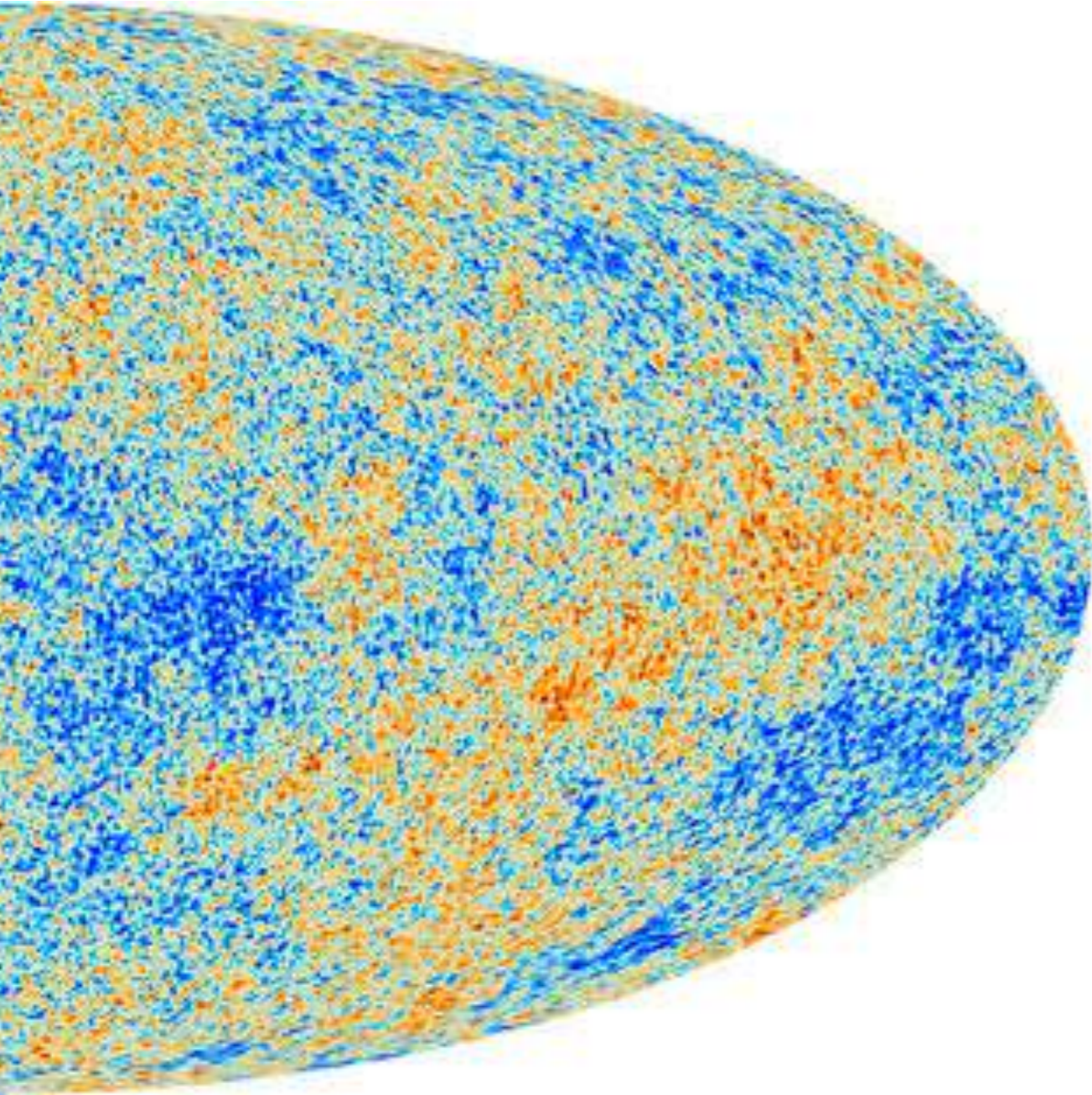
Update from Baryonic Oscillation Spectroscopy Survey



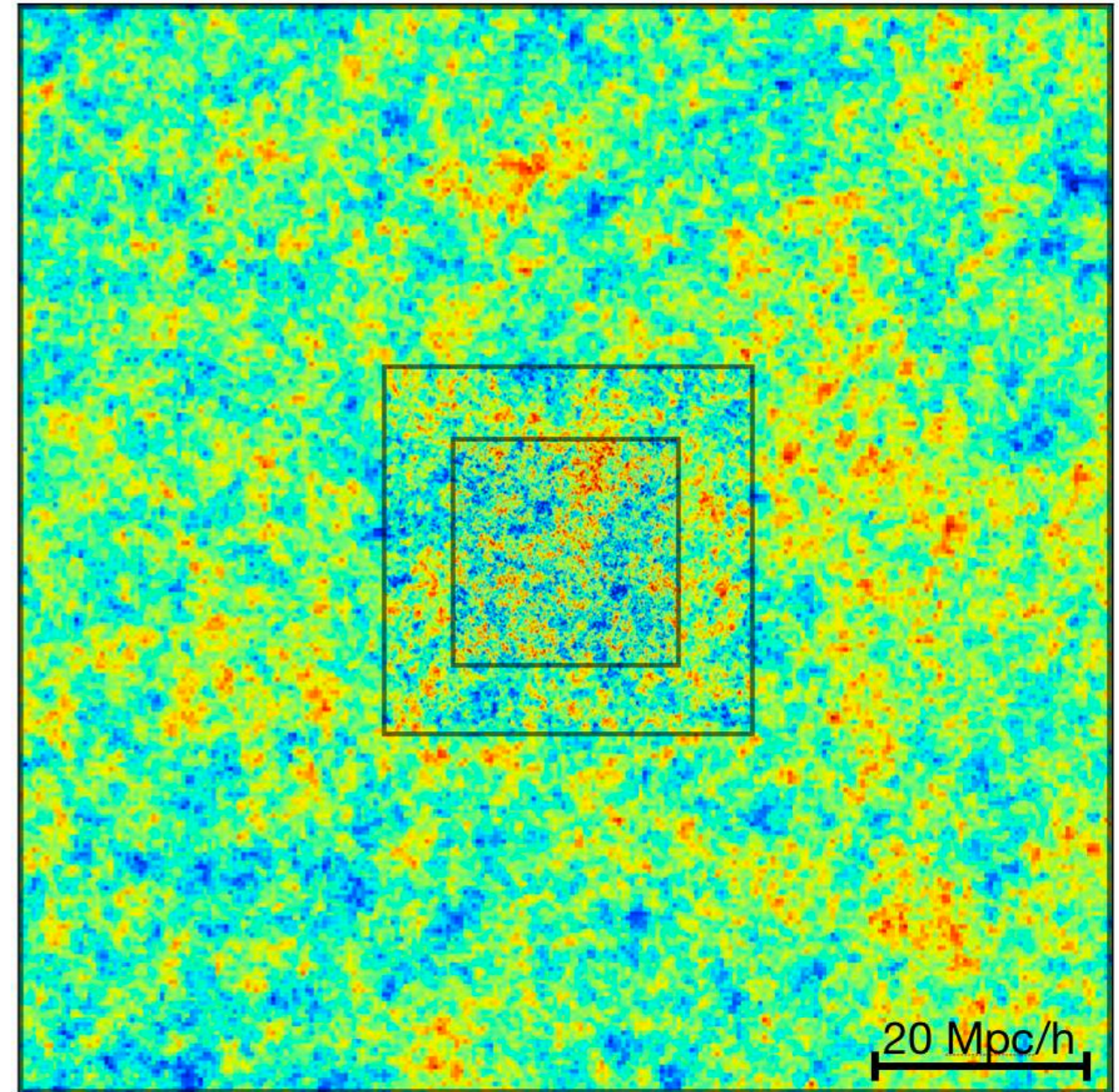
“Fingers” = artifacts of sky coverage

CMB statistics provide initial conditions for structure formation (N -body) simulations

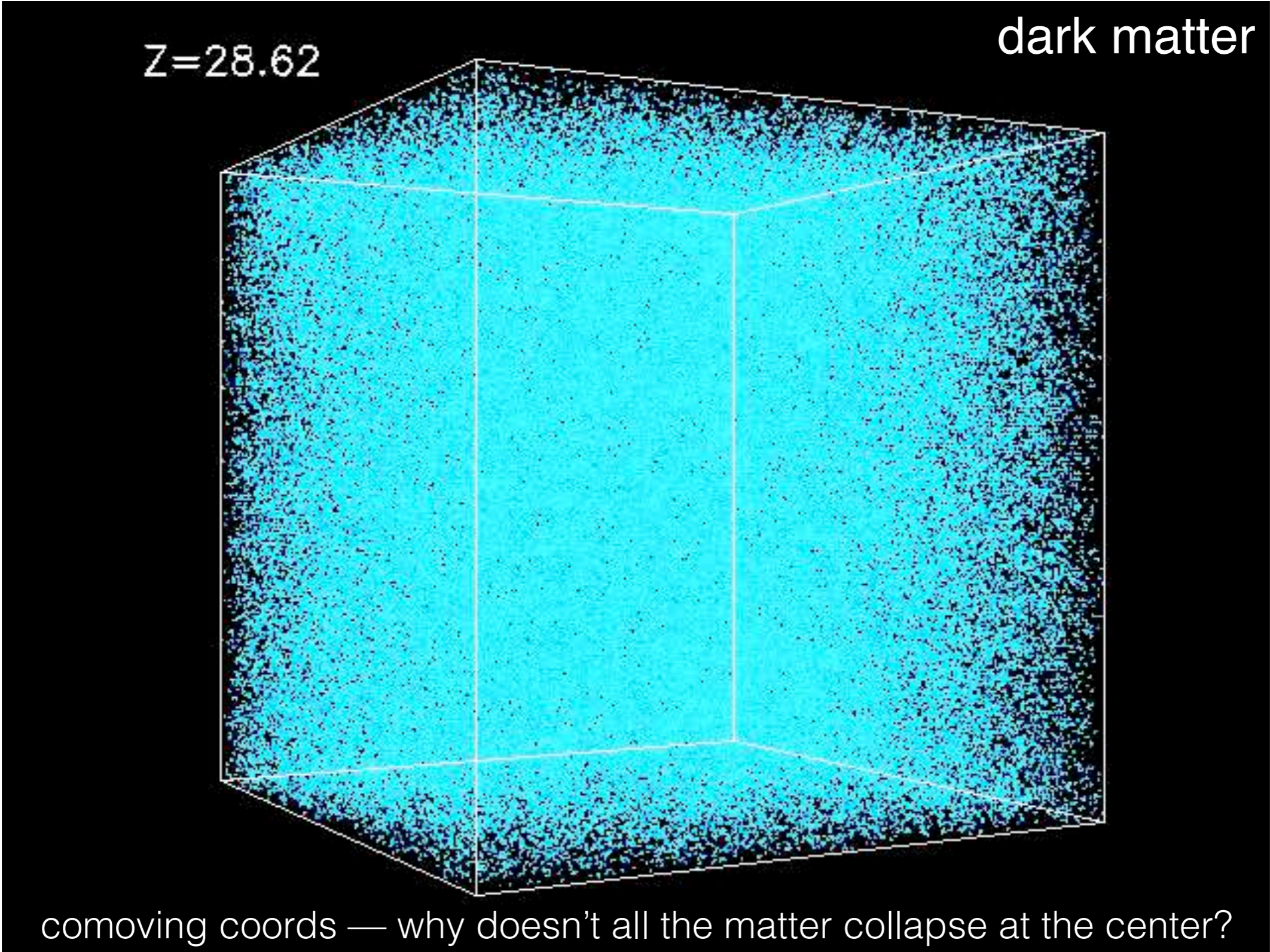
Microwave sky



Simulation ICs

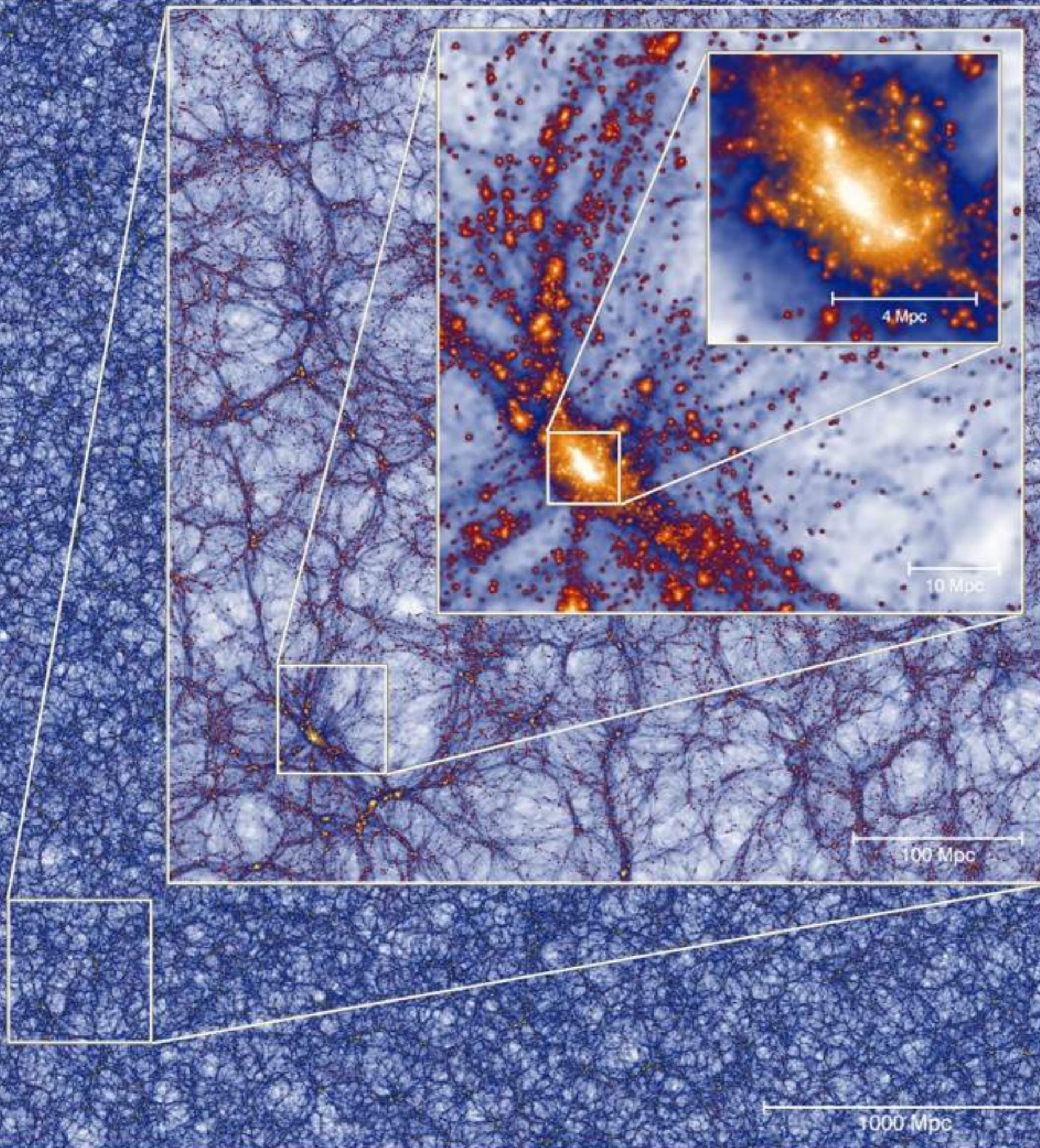


Large-scale structure formation in Λ CDM

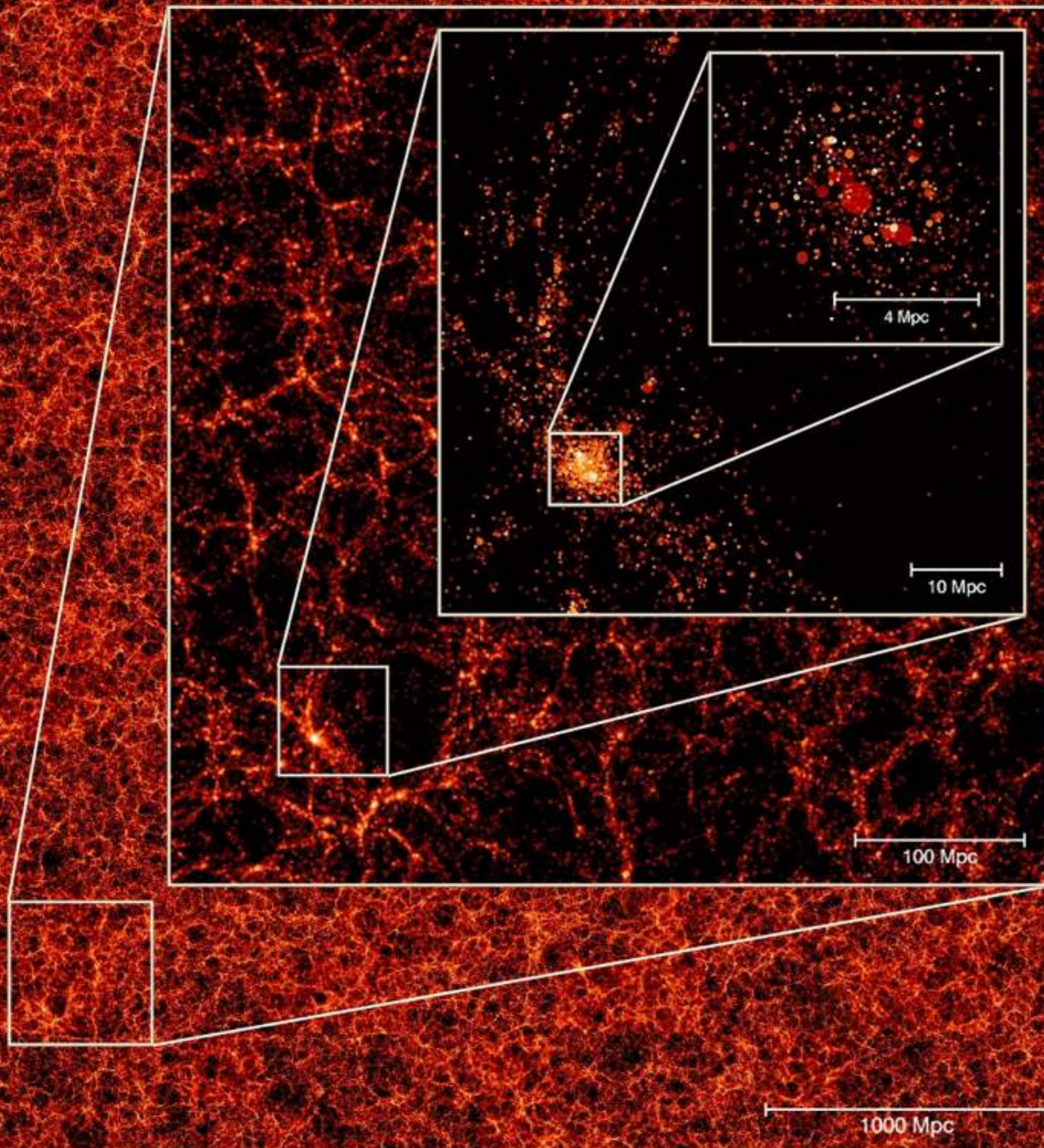


Density peaks (dark matter halos) are the sites of galaxy

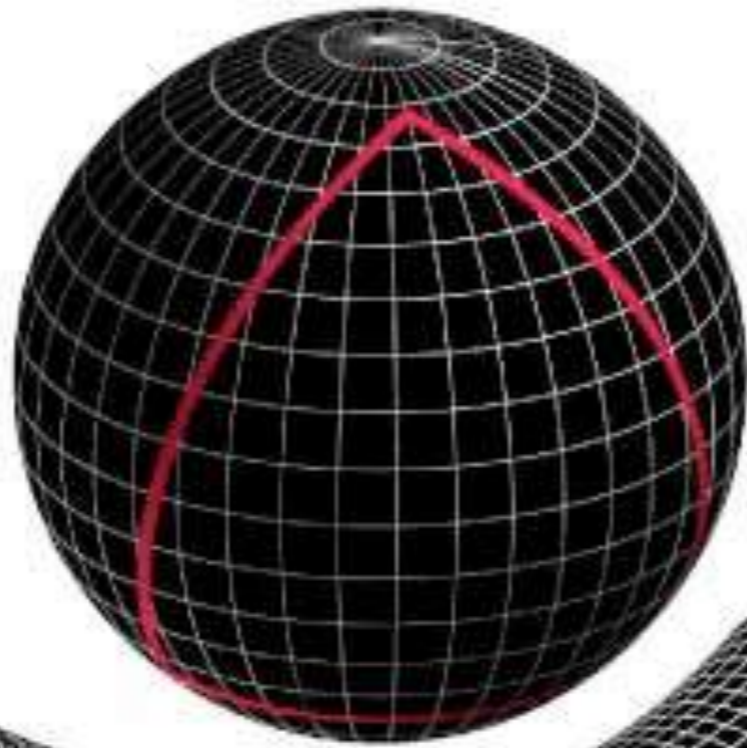
Dark matter
in Millennium-XXL
simulation
(Angulo+12)



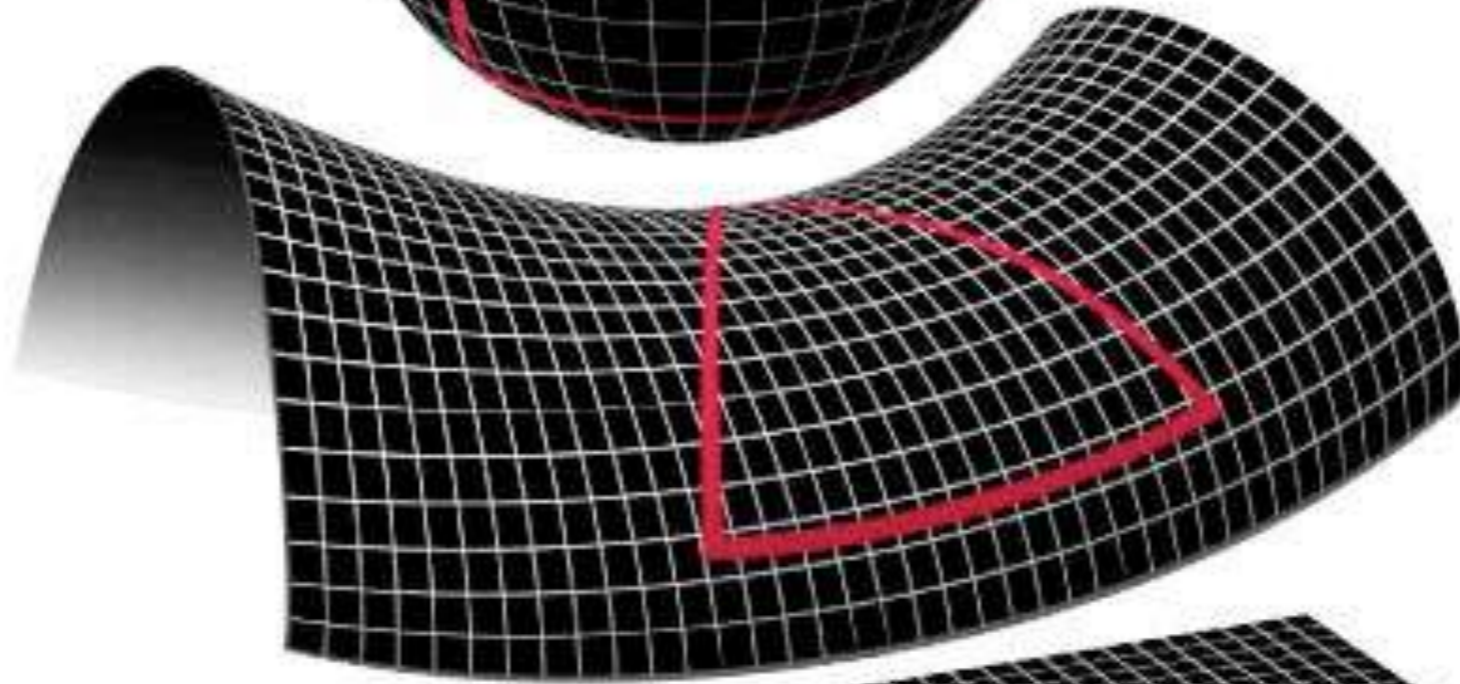
Galaxies
in Millennium-XXL
simulation
(Angulo+12)



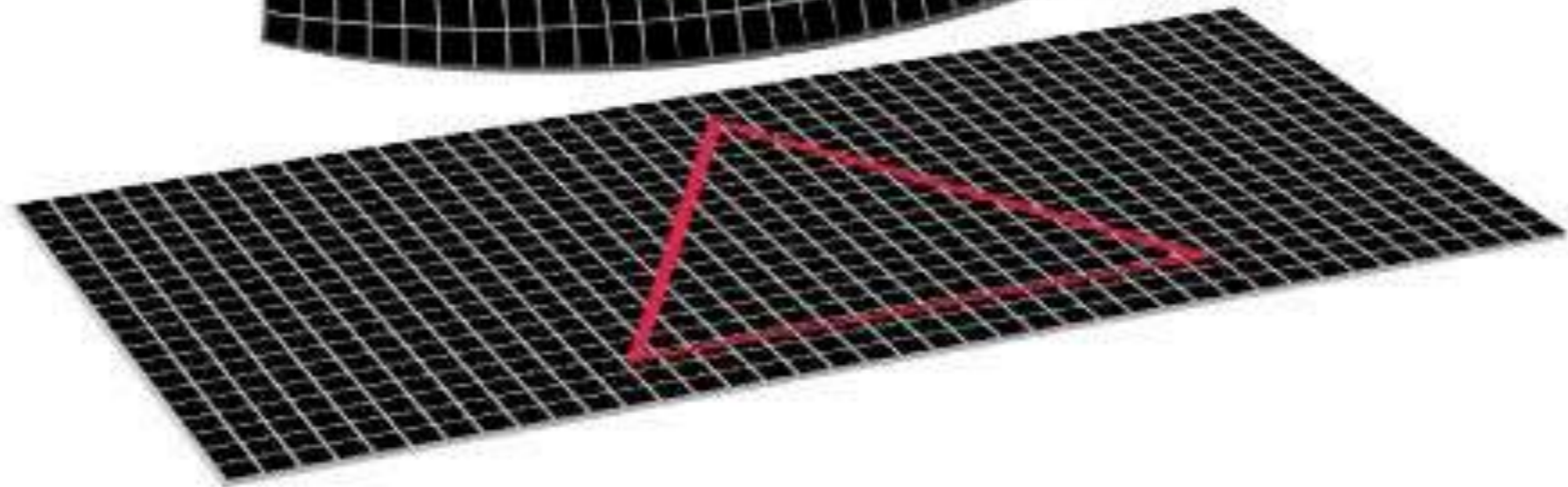
closed



open



flat



Measuring the Universe's geometry using the CMB

